



ARC Centre of Excellence for
Antimatter-Matter
Studies



Fluid Modelling of Plasmas and Swarms: The Big Picture

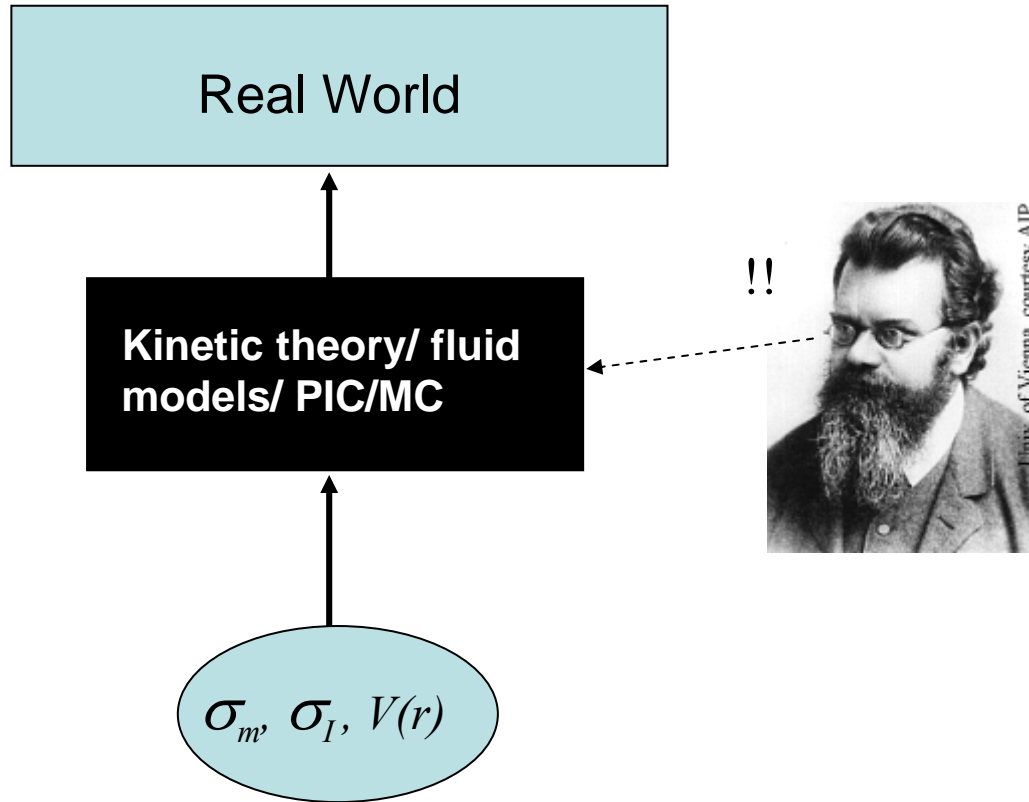
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Acknowledgment

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Microscopic \rightarrow Macroscopic



Kinetic Theory

- Since systems are often far from equilibrium $f \neq$ Maxwellian
- Mean free path analysis too crude to be of much use
- Must solve Boltzmann's equation for the phase space distribution function $f(\mathbf{r}, \mathbf{c}, t)$
- Calculate measurable quantities as velocity "moments", e.g., average velocity

$$\langle \mathbf{c} \rangle = \int \mathbf{c} f(\mathbf{r}, \mathbf{c}, t) d^3\mathbf{c} / \int f(\mathbf{r}, \mathbf{c}, t) d^3\mathbf{c}$$

- Boltzmann's equation plays same role in statistical mechanics as Schrödinger's equation in quantum mechanics, e.g.,

$$\psi(\mathbf{r}, t) + \text{calculation of expectation values } \langle \dots \rangle$$

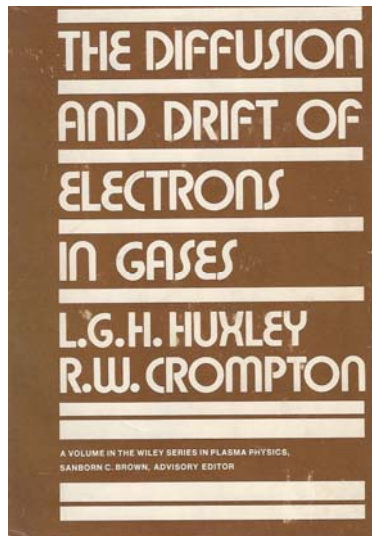
Kinetic Theory of e^\pm , μ^\pm , A^\pm in gases

Modern era

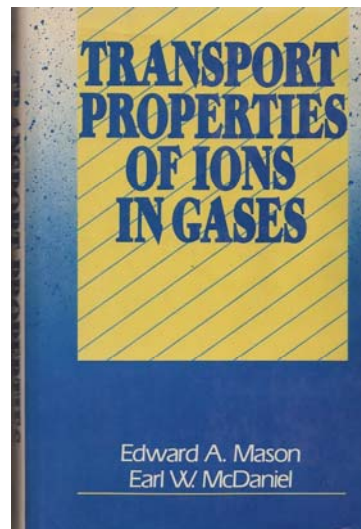
$t > 1975$ (ions), $t > 1979$ (electrons, positrons, muons)

History

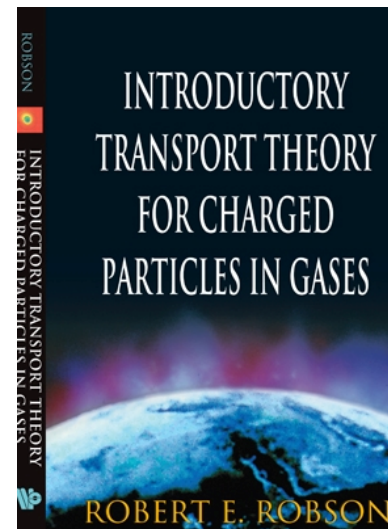
- *Boltzmann 1872* (classical elastic collisions) \rightarrow Fokker-Planck equation
- *Wang-Chang, Uhlenbeck, de Boer 1951* (semi-classical inelastic collisions)
- *Waldmann (1958) - Snider (1961)* (quantum kinetic equation)



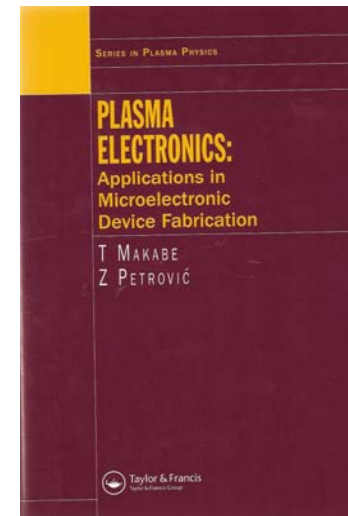
1974



1988



2006



Fluid Modelling

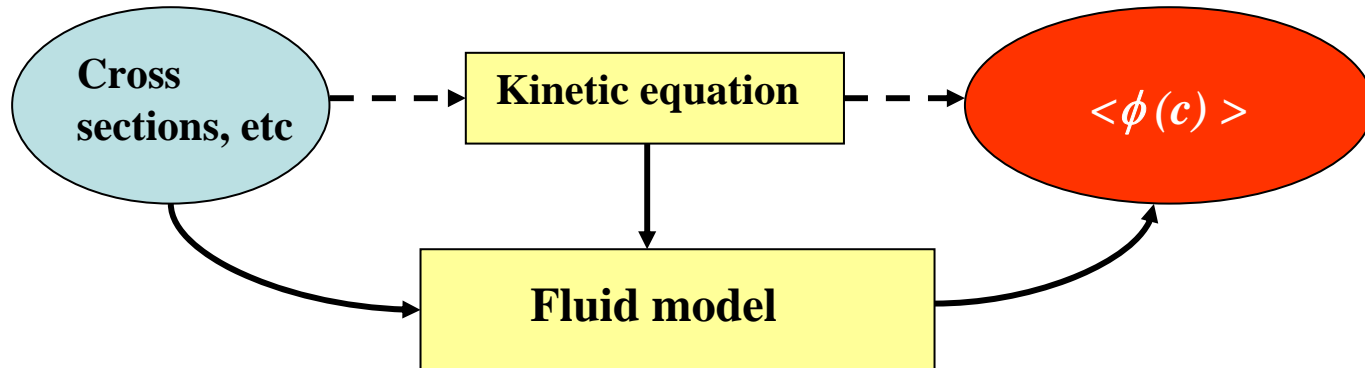
- “Short cut” method \rightarrow $\langle \dots \rangle$ directly **without** solving Boltzmann’s equation - projection of phase space $(\mathbf{r}, \mathbf{c}, t)$ onto configuration space (\mathbf{r}, t)
- Motivation - **computationally economical**
- Fluid equations \equiv adaptation of $\mathbf{F} = m \mathbf{a}$ to fluids + closure *assumptions* to make equations useful
- Numerous *ad hoc*, unphysical assumptions \rightarrow proliferation of fluid models

Assume a...



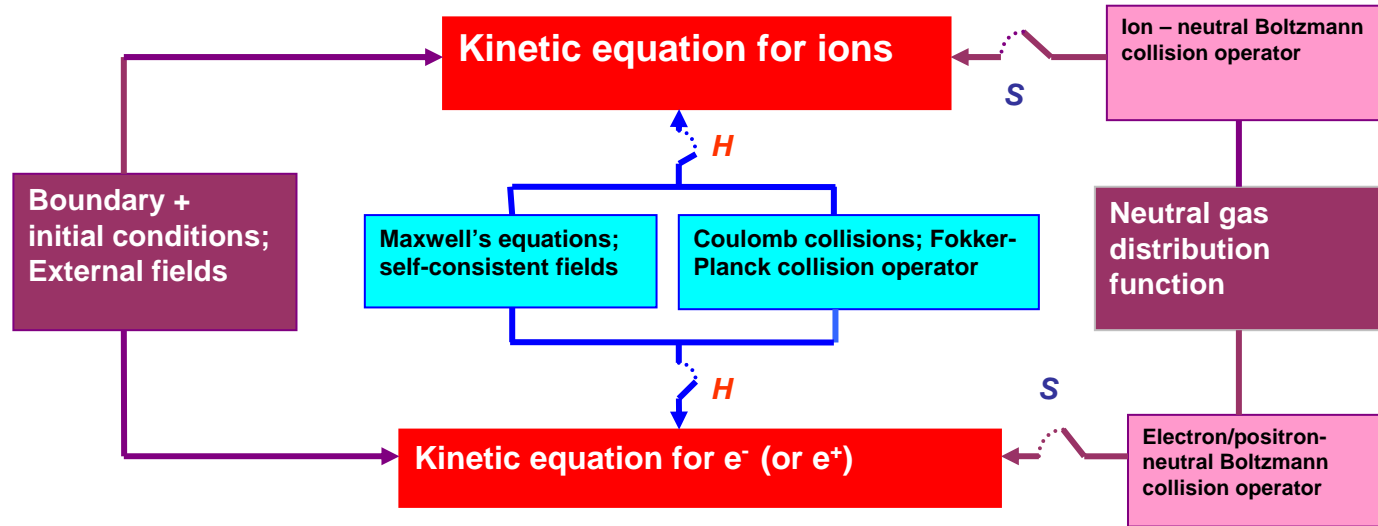
Towards a Physically-based Fluid Model

Aim: To develop physically-based equations valid for *all* types of charged species (e^\pm , A^\pm , ...)



- Form balance equations $\int dc \phi_i(c) \times \text{Kinetic equation}$
 - Systematic approximation required to \rightarrow useful equations
 - Benchmarking required to establish accuracy
 - “One size fits all” possible? After all, there is only one $F = ma$
- Robson, White and Petrović, Rev. Mod. Phys. (2005)

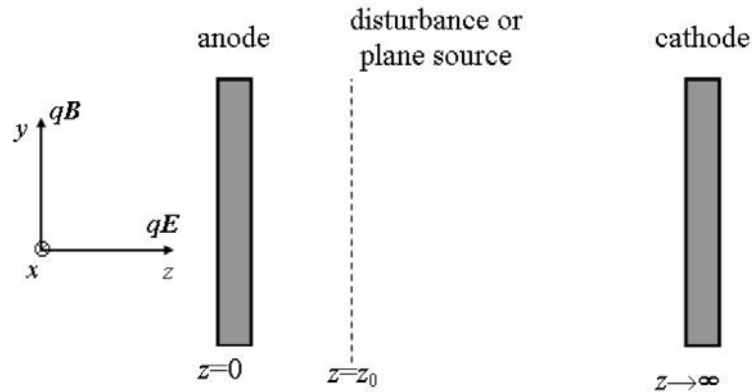
Swarm and plasma modelling



- *Swarm* (= “test particle”) limit - “switches” **H** are open, while the “switches” **S** are closed. Collective effects are usually absent in this limit ($\lambda_D > L$)
- *Hot, fully ionized plasmas* - “switches” **H** are closed while the “switches” **S** are open. Coulomb interaction and collective effects (waves, instabilities) dominate ($\lambda_D \ll L$)
- *Low temperature, collision-dominated plasmas* are intermediate between these two limits, and generally *all* factors must be taken into consideration, i.e., all “switches” **S** + **H** are closed

Benchmark Model

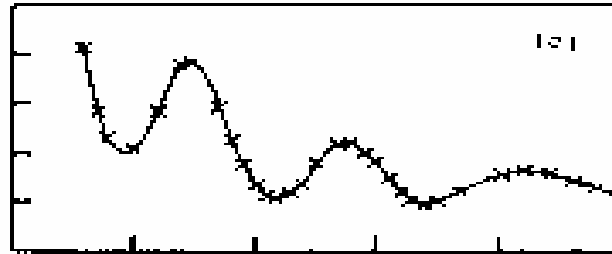
Focus on *relaxation phenomena* in plane-parallel geometry
(G. Petrov et al 1996; B. Li *et al*, J Phys B 2000; Phys Rev E 2006)



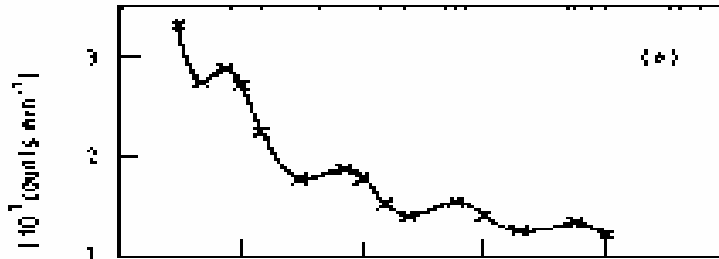
- Electrons emitted at a steady rate from source into infinite gas
- Swarm limit, external fields, no ionisation or attachment
- Diffusion equation \rightarrow unphysical results (both ions and electrons)
- Only a carefully formulated fluid model \rightarrow physically meaningful results
- Exact analytical solution of Boltzmann equation known for $v_m = \text{constant}$
- Solution of Boltzmann equation for real gases - accuracy $\sim 2\%$

Fletcher J. Phys. D(1985) (photon flux technique)

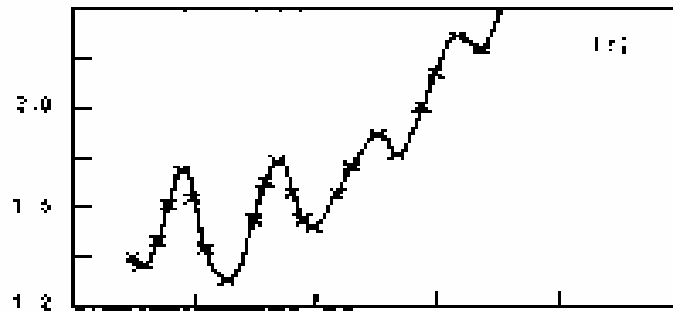
Intensity
(10^3 counts/min)



Ne



Ar

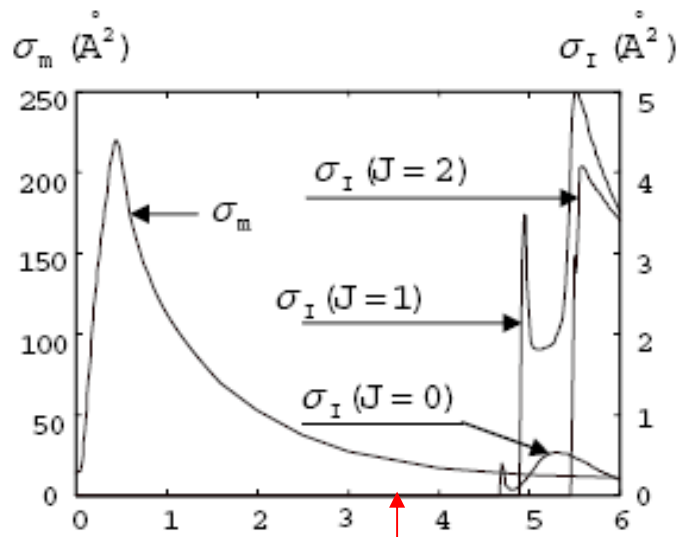


He

z (cm)

Oscillations occur only in a 'window' of voltages and gas pressures

(e⁻, Hg) Franck-Hertz oscillations

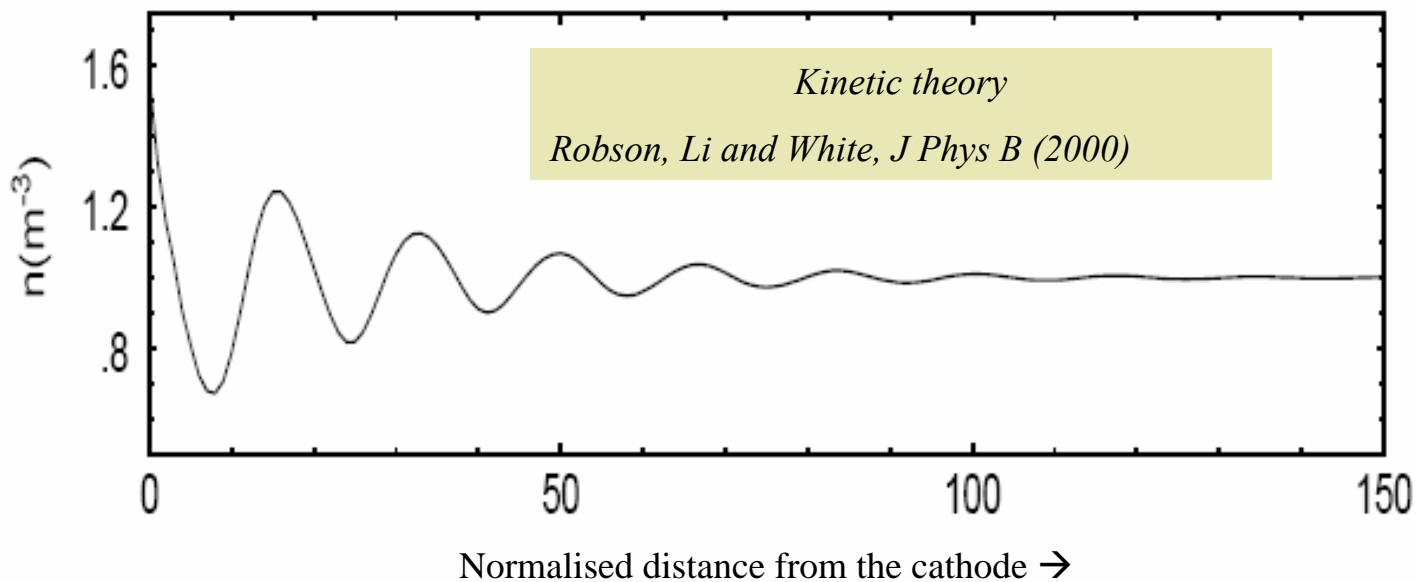


Cross sections

Hanne, Amer. J. Phys. (1988)

England and Elford, Aust. J. Phys. (1991)

Ps formation threshold



Fluid Equations

$$\frac{\partial \Gamma}{\partial z} = 0$$

$$\frac{2}{3} \frac{\partial(n\varepsilon)}{\partial z} = neE - nmv_m(\varepsilon)v$$



$$-\frac{1}{v_e} \left[v \frac{\partial \varepsilon}{\partial z} + \frac{2\varepsilon}{3} \frac{\partial v}{\partial z} + \frac{1}{n} \frac{\partial J_q}{\partial z} \right] = \varepsilon - \frac{1}{2} M v^2 + \Omega(\varepsilon)$$

Inelastic collision term $\Omega = \sum_I \epsilon_I (\vec{v}_I - \overleftarrow{v}_I) / v_e(\varepsilon)$

To *close* the equations, need to express heat flux J_q in terms of n, v, ε

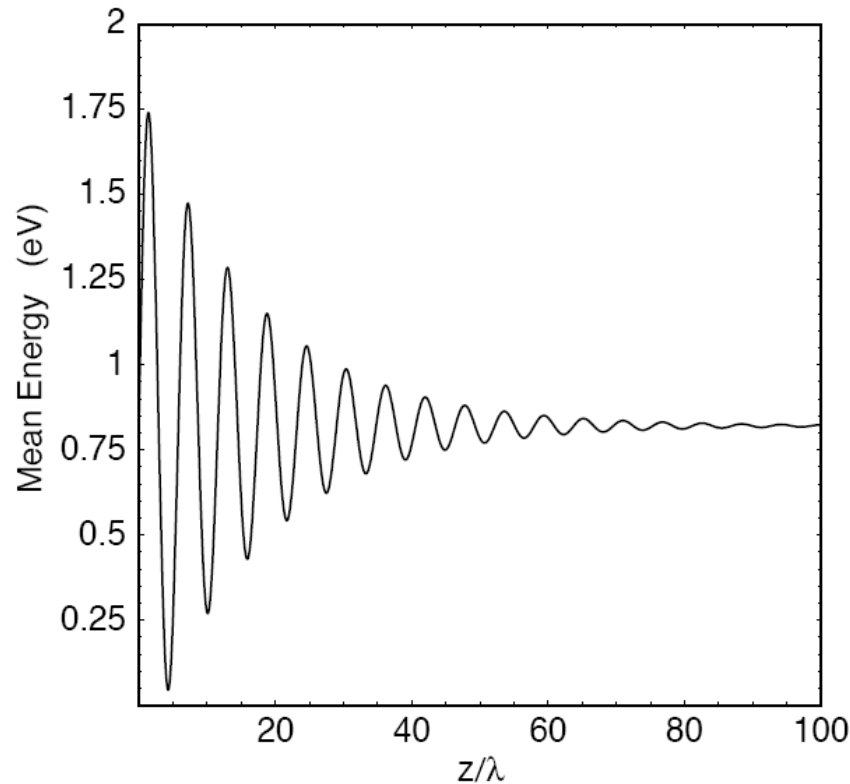
Ansatz $J_q = -\frac{2}{3m} \frac{\partial}{\partial z} \left[\frac{n\xi(\varepsilon)}{v_m(\varepsilon)} \right] + \frac{(5-2p)}{3} \frac{na\varepsilon}{v_m(\varepsilon)} - \frac{5}{3} \Gamma \varepsilon \quad \left(p = \frac{d \ln v_m}{d \ln \varepsilon} \right)$

Expression *exact* for $v_m(\varepsilon) \sim \varepsilon$ and then $\xi(\varepsilon) = \varepsilon^2$

- Fourier ansatz $J_q = -\lambda \partial \varepsilon / \partial z$ 
- $J_q = 0$ 

Surendra and Dalvie (1993)

Mean energy in the “window”



Mean electron energy at $E/N = 6.5 \text{ Td}$ as a function of normalized distance downstream from the source. These Franck-Hertz oscillations are characteristic of all physical properties in the “window” region of E/N - mean velocity, number density, heat flux, although there are phase differences

References

Kinetic modelling

'Advances in low temperature r.f. plasmas' , Ed. T. Makabe (2002)

Robson, White and Morrison, J Phys B 36, 4127 (2003)

Porokhova et al, Phys Rev E 71, 066406 and 066407 (2005)

Pinhao et al, PSST 13, 719 (2005)

Petrović et al, PSST 16, S1-12 (2007)

Fluid Modelling

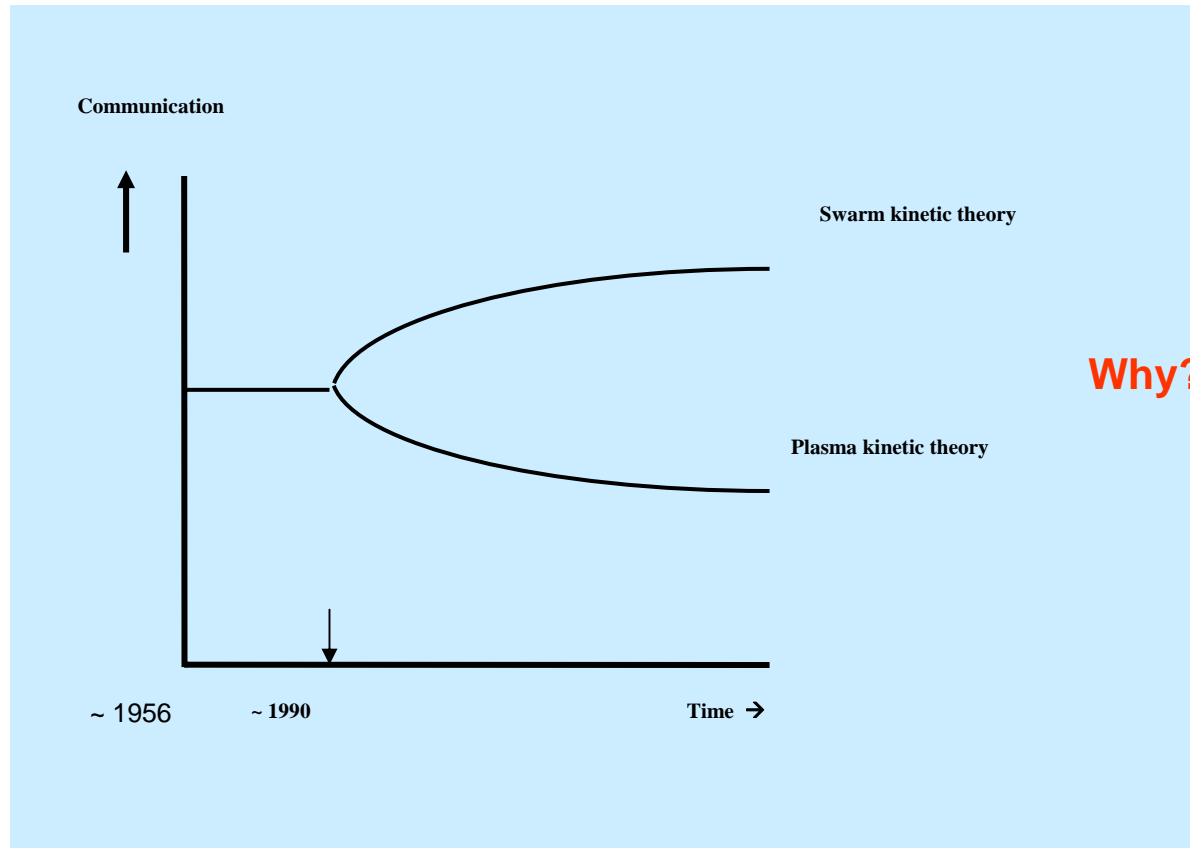
Hagelaar and Pitchford, PSST 14, 722 (2005)

Robson et al, Rev. Mod. Phys. 77, 1303 (2005)

Viehland and Goeringer, J. Phys. B 38, 3987 (2005)

Viehland, GEC Invited Talk (2006)

Swarm vs Plasma Literature



Asymptotic Regime

Linearize equations far downstream from source: $n(z) = n_\infty + n_1 e^{Kz}$ etc.

Three equations \rightarrow cubic secular equation, dimensionless “wave number” $K = \frac{3}{2} \frac{eE}{\varepsilon_\infty} \kappa$

$$-\frac{3}{2}\alpha(\bar{p} - p - 1)\kappa^3 + \left(\frac{3}{2}\alpha\bar{p} - \frac{5}{2}p + p^2\right)\kappa^2 + \left(\gamma + 2p - \frac{7}{2}\right)\kappa - (\gamma + 2p) = 0$$

$$\bar{p} = \frac{\varepsilon_\infty \xi'(\varepsilon_\infty)}{\xi(\varepsilon_\infty)}$$

$$\alpha = \frac{\xi(\varepsilon_\infty)}{\varepsilon_\infty^2}$$

Benchmark model

$M = 4.0 \text{ a.m.u.}, T_g = 0$

$\sigma_m = 6 A^2$

$\sigma_l = 0.1 A^2$

$0 \quad \varepsilon_l = 2 \text{ eV} \quad \varepsilon$

Key parameter

$$\gamma = \frac{(1 + \Omega')\varepsilon_\infty}{\varepsilon_\infty + \Omega}$$

